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# Scheelite-Type Rare Earth Molybdates(V) and  $(VI)$ . Eu<sub>x</sub>MoO<sub>4</sub> and LnMoO<sub>4</sub>

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New tetragonal scheelite-type compounds of stoichiometry  $Eu^{II}_{x}$ MoO<sub>4</sub> (0.67  $\le x \le 1.00$ ) and  $Ln^{III}$ MoO<sub>4</sub> (Ln = Y, Gd-Lu) have been prepared by heating reaction mixtures in sealed tubes. The series  $\rm Eu_xMoO_4$  changes from  $\rm Eu^{II}MoV^O_4$  to  $Eu^{III}$ <sub>0,67</sub>Mo<sup>V1</sup>O<sub>4</sub> with evidence from ir spectra and lattice parameter shifts of distortion and multiple site symmetry of the MOO, *2-* tetrahedra. Density measurements indicate that the solid solutions are characterized by Eu lattice vacancies. The  $Eu_{0.67}$ MoO<sub>4</sub> is readily converted to the monoclinic Eu<sub>2</sub>(MoO<sub>4</sub>), structure, an ordered vacancy modification of the scheelite structure. Infrared spectra and lattice parameter shifts of the LnMoO<sub>4</sub> phases are compatible with the distribution Ln<sup>111</sup>-Mo<sup>V</sup>O<sub>4</sub>. Magnetic susceptibility measurements on YMoO<sub>4</sub> and GdMoO<sub>4</sub> show an effective moment of 1.31  $\mu_B$  for Mo<sup>V</sup>. This is reduced from the spin-only value of 1.73  $\mu_B$  for d<sup>1</sup> by spin-orbit coupling. These are the first compounds reported in which all the Mo is present as  $Mo<sup>V</sup>O<sub>4</sub>$  tetrahedra.

# Introduction

Preparations of mixed oxides of the rare earths and group VI transition elements nominally in the  $5+$  oxidation state have led to strikingly different results in the cases of chromium and tungsten. In the latter case, the products are tungsten bronzes  $\text{Ln}_x \text{WO}_3$   $(0 \lt x \lt 0.33)^1$  while chromium forms compounds of huttonite or zircon structures, $<sup>2</sup>$  with composi-</sup> tion  $LnCrO<sub>4</sub>$ , where the chromium is present in the form of distorted Cr043- tetrahedra with a chromium valence of *5+.*  These compounds are isostructural with the corresponding phosphates, arsenates, and vanadates.<sup>3</sup> The arsenates and vanadates are reported to transform at high pressure to a scheelite-type phase which can be quenched and retained metastably at 1 atm.<sup>4</sup> Stubican has found that rare earth niobates,  $LnNbO<sub>4</sub>$ , and tantalates,  $LnTaO<sub>4</sub>$ , transform to a high-temperature nonquenchable scheelite-type phase.<sup>5</sup> At the time this research was undertaken, there were no reports of reduced molybdenum rare earth oxides, nor was there any report of compounds containing pentavalent molybdenum in  $MoO<sub>4</sub><sup>3-</sup>$  anions. Recently a report of Na<sub>3</sub>MoO<sub>4</sub> has appeared,<sup>6</sup> but its structure is not yet known, although the authors suggest a structure derived from the rock salt structure. In a dissertation by Hubert,<sup>7</sup> structures of the type  $Ln_5Mo_3O_{16}$ were reported. The susceptibility measurements were compatible with  $Ln_5Mo<sup>V1</sup><sub>2</sub>Mo<sup>V</sup>O<sub>16</sub>$ . In the present work, we offer evidence that all the molybdenum in the  $LnMoO<sub>4</sub>$  compounds is in the *5+* oxidation state and that, in the defect solid solutions involving europium, it remains in the form of molybdate(V1) anions.

#### Experimental Section

The  $Eu<sub>x</sub>MoO<sub>4</sub>$  compounds were prepared by heating a mixture of the appropriate molar ratios of  $MoO<sub>2</sub>$ ,  $MoO<sub>3</sub>$ , and Eu<sub>2</sub>O<sub>3</sub> in evacuated and sealed silica tubes for 1–2 hr at  $1200^{\circ}$  except for  $\rm Eu_{o.80}MoO$ which was heated for 2 hr at  $1050^\circ$ . The compounds were deep brown-black except for  $Eu_{0.70}MOQ_4$  and  $Eu_{0.67}MOQ_4$  which were gray. McCarthy has reported the preparation of EuMoO<sub>4</sub>.<sup>8</sup> The LnMoO, compounds were prepared by heating thoroughly mixed equimolar ratios of  $MoO<sub>2</sub>$ ,  $MoO<sub>3</sub>$ , and  $Ln<sub>2</sub>O<sub>3</sub>$  at  $1200°$  for 1 hr in evacuated and sealed silica tubes. The compounds are all black. Attempts **at** preparing LaMo0,-SmMoO, were unsuccessful. The prep-

- **(7)** P.-H. Hubert, Thesis, 1'Universite Claude Bernard, **Lyon I,**  France, **1973.** 
	- **(8)** G. **J.** McCarthy,Mater. *Res. Bull., 6,* **31 (1971).**

arations were black. However, the powder patterns were quite similar to those of the starting rare earth oxides  $Ln<sub>2</sub>O<sub>3</sub>$ . The oxides used were Baker and Adamson MoO<sub>3</sub> (99.5% minimum), Alfa Inorganics **MOO,** (99% minimum), and ROC/RIC rare oxides (99.9% minimum).

Lattice parameter measurements for all powders were calculated by a least-squares program from data taken with graphite-monochromated Cu K<sub>a</sub> radiation using W or Si as internal standards. The infrared measurements were taken **on** a Perkin-Elmer Model **521** spectrophotometer. Samples were pressed in KBr pellets with the exception of  $Eu_{0.67}MoO_4$  which was ground in Nujol and placed between KBr disks.  $Eu_{0.67}MoO_4$  would convert to the stable monoclinic Eu<sub>2</sub>- $(MOO<sub>4</sub>)<sub>3</sub>$  phase when pressed in KBr pellets. Density measurements of  $Eu_{0.90}MoO_4$ ,  $Eu_{0.80}MoO_4$ , and  $GdMoO_4$  were measured by displacement of water. The magnetic susceptibility measurements of  $\text{YMoO}_4$  and  $\text{GdMoO}_4$  were obtained on a vibrating-sample magnetometer from room temperature to liquid nitrogen temperature.

### Lattice Parameter Measurements

Lattice parameters for the LnMoO<sub>4</sub> and Eu<sub>x</sub>MoO<sub>4</sub> compounds are shown in Tables I and 11, respectively. **A** plot of  $V^{1/3}$  *vs. r* for Ln<sup>3+</sup> eight-coordinate rare earth radii<sup>9</sup> is linear within experimental error except for Ho<sup>3+</sup> and Y<sup>3+</sup> (Figure 1). The poor fit at  $Ho^{3+}$  and  $Y^{3+}$  is most likely due to uncertainties in the reported ionic radii. Greis and Petzel<sup>9</sup> calculated ionic radii for **v111Y3+** and v1nHo3+ as 1.004 and 1 .OOO **A,** respectively. However, extensive data on rare earth Al, Ga, and Fe garnets<sup>10</sup> show differences in lattice parameters of no more than one part per thousand for **Y3'**  and Ho<sup>3+</sup> garnets, indicating that the ionic radii of  $VIIIY^{3+}$ and **V111H03+** are closer than the difference indicated by Greis and Petzel for the trifluorides. A plot of *V vs.* **x** for Eu,Mo04 drawn in Figure *2* shows more deviation from a straight line than the plot of Figure 1. This is probably due to the uncertainty in the stoichiometry of europium. The rare earth molybdates *(Y,* Gd-Lu) do not exhibit this nonstoichiometry. Attempts to synthesize  $Y_x$ MoO<sub>4</sub> and Gd<sub>x</sub>-MoO<sub>4</sub> where  $x < 1.00$  resulted in the stoichiometric phase plus the expected impurities. Density measurements on  $Eu_{0.80}MoO_4$  and GdMoO<sub>4</sub>, shown in Table III, agree within experimental error with the X-ray density. The low densities of  $Eu_{0.90}$ MoO<sub>4</sub> and YMoO<sub>4</sub> are most likely due to voids between crystallites. Both compounds are soft sinters at the temperature of preparation. This indicates that the assumed stoichiometry of the  $Eu<sub>x</sub>MoO<sub>4</sub>$  compounds is correct and can be formulated as  $Eu_{x}\Box_{1-x}MoO_{4}$ .

# Infrared Measurements

Infrared spectra of  $EuMoO<sub>4</sub>$  and  $SrMoO<sub>4</sub>$  shown in

**(9) 0.** Greis and T. Petzel, *2. Anorg. Allg. Chem.,* **403, 1 (1974). (IO)** M. Kokta, Thesis, Department of Chemical Engineering, Newark College **of** Engineering, Newark, N. **J., 1972.** 

**<sup>(1)</sup>** W. Ostertag,Inorg. *Chem.,* **5, 758 (1966).** 

**<sup>(2)</sup>** H. Schwarz, *2. Anorg. Allg. Chem.,* **323,275 (1963).** 

**<sup>(3)</sup> H.** Schwarz, *Z. Anorg. Allg. Chem.,* **323, 33 (1963). (4) V. S.** Stubican and R. Roy, *Z. Kristallogr., Kristallgeometrie,*  **(5) V. S.** Stubican, *J. Amer. Ceram.* **SOC.,** *41,* **55 (1964).**  *Kvistallphys., Kristallchem,* **119, 90 (1963).** 

*<sup>(6)</sup>* H. **Kessler,** C. Ringenbach, L. Lavielle, and **A.** Hatterer, *C. R. Acad. Sci., Ser. C.,* **271, 763 (1973).** 

<b>Table I.</b> Lattice Parameters for Tetragonal LnMoO <sub>4</sub> <sup>a</sup>					
	a, A	c, A	$V$ . A $^3$		
$YMO_{4}$	5.170(1)	11,000(4)	294.1(2)		
EuMoO <sub>a</sub>	5.392(2)	11.992(5)	348.7(3)		
$Eu_{0.95}Gd_{0.95}MoO4$	5.374 (2)	11.942 (6)	344.9 (2)		
$Eu_{0.85}Gd_{0.15}MoO4$	5.351 (2)	11.850(6)	339.4 (2)		
$Gd_{0.85}Eu_{0.15}MoO_4$	5.220(1)	11.170 (2)	304.4(1)		
GdMoO <sub>a</sub>	5.214(2)	11.180 (6)	304.0(3)		
$TbM_0Oa$	5.202(1)	11.110 (2)	300.7(1)		
DyMoO <sub>A</sub>	5.187(1)	11.064(4)	297.7(1)		
$H_0MO_A$	5.175(1)	11.008 (4)	294.9(1)		
ErMoO,	5.156(1)	10.966 (4)	291.5(2)		

*a* Least-squares deviations in parentheses.

Table II. Lattice Parameters for Tetragonal  $Eu<sub>x</sub>MoO<sub>a</sub><sup>a</sup>$ 

	a, A	c. A	$V$ . A <sup>3</sup>	
EuMoO <sub>a</sub>	5.392(2)	11.992(5)	348.7(3)	
$Eu_{0.90}MoO4$	5.343(3)	11.821(5)	337.5(3)	
$Eu_{0.80}MoO4$	5.299(2)	11.769(4)	330.6(2)	
$Eu_{0.20}MoO4$	5.255(4)	11.608(9)	320.5(5)	
$Eu_{0.67}MoO4$	5.232(2)	11.528(6)	315.6(3)	

 $\frac{TmMO_{4}}{YbMoO_{4}}$  5.141 (1) 10.962 (3) 288.6 (1)<br> $\frac{5.127 (1)}{10.880 (2)}$  286.0 (1) YbMoO<sub>4</sub> 5.127 (1) 10.880 (2) 286.0 (1)<br>LuMoO<sub>4</sub> 5.117 (1) 10.857 (3) 284.3 (1)

 $10.857(3)$ 

*a* Least-squares deviations in parentheses.





Figure 1.  $V^{1/3}$  vs. eight-coordinate radii of Ln<sup>3+</sup> ions<sup>9</sup> for LnMoO<sub>4</sub> compounds.

Figure 3 are very similar, as would be expected for  $MoO<sub>4</sub><sup>2</sup>$ tetrahedra. The spectra of the europium-deficient  $Eu<sub>x</sub>MoO<sub>4</sub>$ compounds show a distortion of the  $MoO<sub>4</sub><sup>2</sup>$  tetrahedra. The band at 705 cm<sup>-1</sup> increases in relative intensity with decreasing europium content. This is compatible with multiple-site symmetry for the  $MoO<sub>4</sub><sup>2</sup>$  tetrahedra.

The spectrum of GdMoO<sub>4</sub> in Figure 4 is characteristic of all the LnMoO<sub>4</sub> compounds and is also similar to that of Gd-NbO<sub>4</sub>, reported by Komkov<sup>11</sup> to contain tetrahedral NbO<sub>4</sub><sup>3-</sup> groups, which is also shown in Figure 4. This indicates that  $Mo<sup>VO</sup><sub>4</sub><sup>3-</sup> tetrahedra are present. Solid solutions of  $Gd_{1-x}$ .$  $Eu<sub>x</sub>MoO<sub>4</sub>$  and  $Eu<sub>1-x</sub>Gd<sub>x</sub>MoO<sub>4</sub>$  were prepared where the solid solubility limits of  $x$  were both  $\sim$  0.15. The infrared spectra of some of these solid solutions are shown in Figure 4. The





**Figure 2.** Volume *vs.*  $x$  in Eu<sub>x</sub>MoO<sub>4</sub> compounds.

spectrum of  $Gd_{0.85}Eu_{0.15}MoO_4$  remains unchanged from that of GdMoO<sub>4</sub>. This is a clear indication that  $Eu^{3+}$  enters the  $Gd^{3+}$  sites. If  $Eu^{2+}$  were to enter the sites, the presence of  $Mo<sup>VI</sup>O<sub>4</sub><sup>2-</sup>$  would be necessary for charge compensation. No  $MoO<sub>4</sub><sup>2</sup>$  vibrational modes are present in the spectrum. In addition, the cell volume is only slightly increased above that of  $GdMoO<sub>4</sub>: 304.4 A<sup>3</sup> vs. 304.0 A<sup>3</sup>, indicating substitu$ tution by  $Eu^{3+}$ , not  $Eu^{2+}$ . The case with  $Eu_{0.85}Gd_{0.15}MO_{4}$ is more confusing.  $Gd^{3+}$  must enter  $Eu^{2+}$  sites. This is confirmed by the substantial decrease in the cell volume compared with EuMoO<sub>4</sub>: 339.4  $A^3$  vs. 348.7  $A^3$ . The change in the infrared spectrum is most likely due to the presence of distorted  $Mo<sup>VI</sup>O<sub>4</sub><sup>2-</sup>$  and  $Mo<sup>V</sup>O<sub>4</sub><sup>3-</sup>$  tetrahedra. The strongest broad band of **MoV043-** which would be expected at 650  $cm<sup>-1</sup>$  may be hidden under the broad band of a distorted  $Mo<sup>VIO<sub>4</sub><sup>2</sup></sup>$  tetrahedron. The similarity between the spectra of  $Eu_{0.85}Gd_{0.15}MoO_4$  and  $Eu_{0.90}MoO_4$  suggests that there may be some  $Mo<sup>3-</sup>$  present in the latter. However, the infrared measurements cannot resolve this uncertainty.

#### Magnetic Measurements

 $\text{YMoO}_4$  and  $\text{GdMoO}_4$  in the temperature range 77-300°K. The plots of  $1/\chi \times 10^{-2}$  ((emu)<sup>-1</sup>) *vs. T* for YMoO<sub>4</sub> and  $1/\chi$  ((emu)<sup>-1</sup>) *vs. T* for GdMoO<sub>4</sub> are shown in Figures 5 and 6, respectively: for YMoO<sub>4</sub>,  $C_m = 0.213(1)$ ,  $\mu_{eff} = 1.306$  $(4) \mu_B$ ,  $\theta = -62$  (1)<sup>o</sup>K; for GdMoO<sub>4</sub>,  $C_m = 8.175$  (8),  $\mu_{eff} =$ Magnetic susceptibility measurements have been made on  $8.08\bar{6}$  (5)  $\mu_{\text{B}}$ ,  $\theta = 0.3$  (2)<sup>o</sup>K.

If we assume the  $Mo<sup>V</sup>$  contribution to the susceptibility remains unchanged, the GdIII contribution to the Curie constant  $C_{\text{m(Gd)}} = 7.86$ . This compares favorably with the expected value for  $J = \frac{7}{2}$ ,  $C_m = 7.88$ .

The large negative deviation of the effective moment of Mo<sup>V</sup> from the spin-only value of 1.73  $\mu_B$  is undoubtedly caused by spin-orbit coupling of the lowest lying state and the first excited state. The infrared spectrum of GdMo04



Figure 3. Infrared absorption of  $SrMoO<sub>4</sub>$  and  $Eu<sub>x</sub>MoO<sub>4</sub>$  compounds taken in pressed KBr pellets.  $Eu_{0.67}MoO<sub>4</sub>$  was taken in a Nujol mull between KBr disks. An asterisk indicates filter change.

indicates substantial distortion of the MoV04 **3-** tetrahedron. Most likely the lowest **2E** state of an undistorted tetrahedron is split and the energy gap between this and the next excited state may not be large. The ground state is orbitally nondegenerate causing complete quenching of the ground-state orbital momentum contributions to the effective moment. In octahedral  $AMoF_6$  ( $A = Na$ , K, Nb, Cs) the effective moment has been reported to be as low as 1.5  $\mu_{\mathbf{B}}$ .<sup>12</sup> Since  $\Delta$ (tetrahedral) =  $\frac{4}{9}$   $\Delta$ (octahedral) for equivalent ligands, the crystal field splitting  $\Delta$  for these tetrahedral molybdates is quite reduced from the octahedral MoF<sub>6</sub><sup>-</sup> case. In the case of a completely quenched ground-state orbital momentum  $\mu_{\text{eff}} = \mu_{\text{so}} (1 - \alpha \lambda / \Delta),^{13}$  where  $\mu_{\text{so}}$  is the spin-only moment,

**(1958).** 

(1 **3) B. N.** Figgis and J. Lewis, "Modern Coordination Chemistry," J. Lewis and R. G. Wilkins, Ed., Interscience, **London, 1960. (14)** T. M. **Dunn,** *Trans. Faraday Soc., 57,* **1441 (1961).** 



Figure 4. Infrared absorption of  $Eu_{1-x}Gd_xMoO_4$  and  $GdNbO_4$  $(Mono = monoclinic)$ .



**Figure 5.**  $1/\chi_m \times 10^{-2}$  ((emu)<sup>-1</sup>) *vs.*  $T(^{\circ}K)$  for YMoO<sub>4</sub>.

 $\lambda$  is the spin-orbit coupling constant,  $\alpha = 2$  for the d<sup>1</sup> case, (12) G. B. Hargreaves and R. D. Peacock, *J. Chem. Soc.*,  $3776$  and  $\Delta$  is the crystal field splitting parameter. Dunn<sup>14</sup> has



Figure 6.  $1/\chi_{\text{m}}$  ((emu)<sup>-1</sup>) vs. T (°K) for GdMoO<sub>4</sub>.

calculated the spin-orbit coupling constant for the Mo<sup>V</sup> free ion as  $1030 \text{ cm}^{-1}$ . The actual spin-orbit coupling constant for coordinated Mo<sup>V</sup> is most likely reduced only slightly from this value due to the slight expansion of the electron cloud. If  $\mu_{\text{eff}}$  for Mo<sup>V</sup> in these compounds is 1.306 (4)  $\mu_{\text{B}}$ and  $\alpha\lambda/\Delta \approx 0.23$ , this would require  $\Delta \approx 9000 \text{ cm}^{-1}$ .  $\Delta$  might be determined spectroscopically. However, these compounds are black and the spectrum from the infrared through the visible region is flat with the molybdates diluted in KBr pellets. It might be possible to dilute these compounds in a solid solution with an appropriate scheelite host. However no A<sup>III</sup>B<sup>V</sup>O<sub>4</sub> scheelite phases stable at room temperature are known with the exception of the high-pressure phases of the rare earth arsenates and vanadates.<sup>4</sup> It is also not certain that the distortion of the  $BO_4^3$ <sup>-</sup> tetrahedra in these compounds is closely similar to that in the molybdates. It would be easier to dope these molybdates in rare earth niobates Ln- $Nb<sup>V</sup>O<sub>4</sub>$  which crystallize in the monoclinic fergusonite structure. This is a distortion of scheelite.<sup>11</sup> In addition  $NbO<sub>4</sub>$ <sup>3-</sup> has a similar tetrahedral distortion to those of the molybdates as indicated in its infrared spectrum in Figure 4.

# Conclusions

Lattice parameter and density measurements on the solid solutions  $Eu<sub>x</sub>MoO<sub>4</sub>$  indicate that the molybdenum remains in the  $6+$  oxidation state and that the vacancies are compensated by the presence of Eu<sup>3+</sup> resulting in the formula  $E^{uH}_{1-3x}E^{uH}_{u}E^{uH}_{2x}\Box_x MoO_4$ . Although the infrared spectra of these phases appear to support this formulation, with the appearance of broadening and extra peaks indicating the presence of multiple sites, the presence of some  $MoO<sub>4</sub><sup>3</sup>$  cannot be completely ruled out. The increase in absorption in the region of the main peak of  $MoO<sub>4</sub><sup>3-</sup>$  cannot be attributed to that anion, but it is possible that it may be making some contribution in concentrations too low to permit the appearance of a peak at  $650 \text{ cm}^{-1}$ . The poor resolution of these spectra

and the inability to obtain Raman spectra have precluded any attempt at more detailed analysis.

The disordered scheelite structure of  $Eu_{0.67}MoO<sub>4</sub>$  appears to be similar to that of the high-temperature phase which was suggested to exist above  $848°$  by Jeitschko,<sup>15</sup> who reported that single crystals of  $La_2(M_0O_4)_3$  grown by the Czochralski method fractured on cooling through that temperature. This was attributed to a disorder-order transition. Our  $Eu_{0.67}MoO<sub>4</sub>$ , quenched from 1200°, transformed to the ordered structure of  $Eu_2(M_0O_4)_3$  on grinding with KBr. This indicated it was metastable.

The LnMoO<sub>4</sub> compounds clearly contain MoO<sub>4</sub><sup>3-</sup>, with no evidence for  $MoO<sub>4</sub><sup>2</sup>$  in the infrared spectra. The significant decrease in lattice parameters relative to EuMoO<sub>4</sub> and the similarity of the infrared spectra of  $GdNbO<sub>4</sub>$  and  $GdMo-$ **O4** all strongly confirm the formulation. The infrared spectrum of  $GdNbO<sub>4</sub>$  essentially agrees with that of  $YNbO<sub>4</sub>$  previously reported.16 The cell volume of the high-temperature scheelite form of  $GdNbO_4$  has been reported<sup>17</sup> as 306.2  $A^3$ measured at *850"* which is quite close to our value of 304.0  $A<sup>3</sup>$  for GdMoO<sub>4</sub> taken at room temperature. In addition, the isostructural relation of all members of the series, including  $YMoO<sub>4</sub>$ , and the regular decrease in lattice parameters following the lanthanide contraction leave little room for doubt. These compounds appear to be the first in which the  $MoO<sub>4</sub><sup>3-</sup>$  anion has been reported as the only molybdenum species present.

The growth of single crystals of the intermediate  $Eu<sub>x</sub>MoO<sub>4</sub>$ phase would be of considerable interest, especially if ordered low-temperature phases could be prepared at certain composi-. tions. Flux methods would probably be necessary; samples with  $x > 0.67$  melted incongruently. The LnMoO<sub>4</sub> compounds decomposed to mixtures of  $\text{Ln}_2(\text{MoO}_4)_3$ , MoO<sub>2</sub>, and  $Ln<sub>2</sub>O<sub>3</sub>$  at high temperatures, and low-melting flux techniques might be the only possibility for successful growth.

The reduced moment of  $Mo<sup>V</sup>(4d<sup>1</sup>)$  in YMoO<sub>4</sub> and GdMo- $O<sub>4</sub>$  can be attributed to spin-orbit coupling. It would be useful to be able to determine the splitting parameter  $\Delta$ spectroscopically, preferably in a dilute system. The rare earth niobates  $LnNbO<sub>4</sub>$  would be a first choice as the host for these materials.

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**Registry No.** YMoO,, **52322-39-1;** EuMoO,, **22206-17-3;** Gd-MOO,, **52322-34-6;** TbMoO,, **52322-37-9;** DyMoO,, **52322-324;**  HoMoO,, **52322-35-1;** EIMoO,, **52322-33-5;** TmMoO,, **52322-38-0;**  YbMoO, **,523 22-404** ; LuMoO, **,5 23 22-36-8** ; Eu,MoO,, **5 2322-4 1-5.** 

**<sup>(15)</sup> W. Jeitschko,** *Acta Crystallogr., Sect. B,* **29, 2074 (1973).** 

**<sup>(16)</sup> G. Blasse,** *J. Solid State Chem.,* **7, 169 (1973).** 

**<sup>(17)</sup> K. A. Gingerich and H. E. Blair,** *Advan. X-Ray* **Anal.,** *7,*  **22 (1964).**